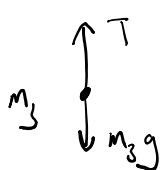
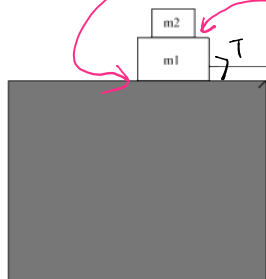
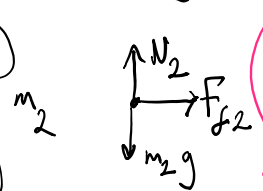


QUIZ-9

Three masses  $m_1=1\text{kg}$ ,  $m_2=1\text{kg}$  and  $m_3=2\text{kg}$  are arranged as shown in the figure. If the friction coefficient between  $m_1$  and  $m_2$  is  $\mu_1=0.1$ , and the friction coefficient between  $m_1$  and the table is  $\mu_2=0.65$ , find the acceleration of all masses and the tension of the cable. (Gravitational acceleration is  $g=10\text{ m/s}^2$ )

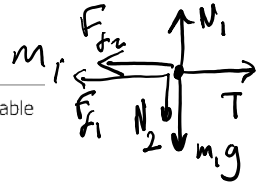
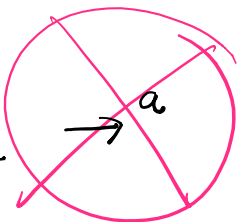


$$m_3 a = m_3 g - T$$



$$N_2 = m_2 g$$

$$m_2 a = F_{f2}$$



$$N_1 = N_2 + m_1 g$$

$$N_1 = (m_1 + m_2) g$$

$$m_1 a = T - F_{f1} - F_{f2}$$

kinetic friction

$$F_{f1} = \mu_2 N_1$$

$$= \mu_2 (m_1 + m_2) g$$

Acceleration of $m_1$	Acceleration of $m_2$	Tension of the cable
$2\text{ m/s}^2$	$1\text{ m/s}^2$	$14$

1° Assume  $m_2$  remains stationary w.r.t.  $m_1$

$$m_3 a = m_3 g - T$$

$$m_2 a = F_{f2}$$

$$m_1 a = T - F_{f1} - F_{f2}$$

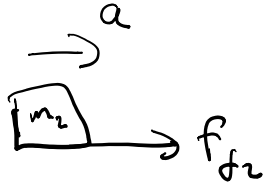
$$(m_1 + m_2 + m_3) a = m_3 g - F_{f1}$$

$$(m_1 + m_2 + m_3) a = m_3 g - \mu_2 (m_1 + m_2) g$$

$$a = \frac{m_3 - \mu_2 (m_1 + m_2)}{(m_1 + m_2 + m_3)} g = \frac{2 - (0.65) 2}{4} 10^5$$

$$= (0.35) \times 5 = 1.75\text{ m/s}^2$$

Assume



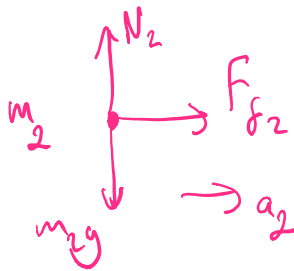
$$F_{f2} = m_2 a = 1.75\text{ N}$$

$$|F_{f2}| \leq \mu_1 N_2 = 0.1 m_2 g = 1\text{ N}$$

$$|F_{f2}| \leq \mu_1 N_2 = 0.1 m_2 g = \underline{1 \text{ N}}$$

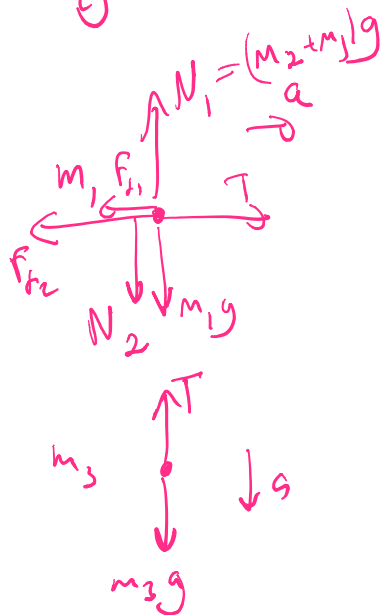
~~$$1.75 \text{ N} \leq 1 \text{ N}$$~~

$m_2$  is not static with respect to  $m_1$ !



$$|F_{f2}| = \mu_1 N_2 = 0.1 m_2 g = \underline{1 \text{ N}}$$

$$m_2 a_2 = F_{f2} \Rightarrow \boxed{a_2 = 1 \text{ m/s}^2}$$



$$T - F_{f1} - F_{f2} = m_1 a$$

$$m_3 g - T = m_3 a$$

$$(m_1 + m_3) a = m_3 g - F_{f1} - F_{f2}$$

$$m_3 g - m_3 a = T$$

$$2(10 - 2) = \sqrt{160}$$

$$3a = 20 - 13 - 1$$

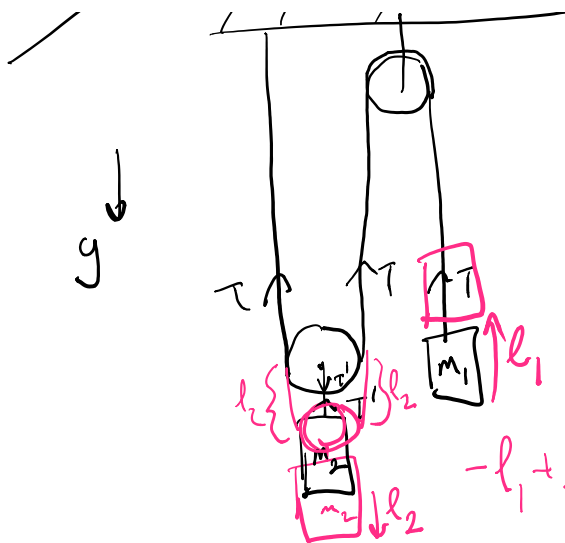
$$a = \frac{6}{3} = \boxed{2 \text{ m/s}^2}$$

$$\begin{aligned} F_{f1} &= N_1 \mu_2 \\ &= 0.65 \cdot 2 \cdot 10 \\ F_{f1} &= 13 \text{ N} \end{aligned}$$

Ex



No friction



- No friction
- inextensible string
- massless pulleys.

Find the acceleration of both masses and the tension in the rope.

$$-l_1 + 2l_2 = 0$$

$$\frac{d}{dt} \rightarrow$$

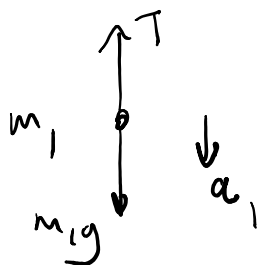
$$-v_1 + 2v_2 = 0$$

$$-a_1 + 2a_2 = 0$$

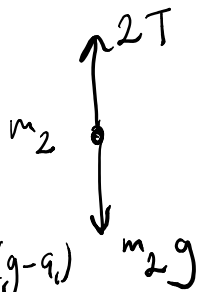
$$a_2 \uparrow$$

$$2a_2 = a_1$$

(3)



$$T = m_1(g - a_1)$$



$$m_2 a_2 = 2T - m_2 g$$

$$m_1 g - T = m_1 a_1$$

$$2(m_1 g - T) = (m_1, 2a_2) 2 \quad \text{①, ③}$$

$$2T - m_2 g = m_2 a_2$$

+

$$2m_1 g - 2T + 2T - m_2 g = 4m_1 a_2 + m_2 a_2$$

$$(2m_1 - m_2)g = (4m_1 + m_2)a_2$$

$$a_2 = \frac{(2m_1 - m_2)}{(4m_1 + m_2)} g$$

$$a_1 = 2 \frac{(2m_1 - m_2)}{(4m_1 + m_2)} g$$

$$T = m_1(g - a_1) = m_1 g \left(1 - \frac{4m_1 - 2m_2}{4m_1 + m_2}\right)$$

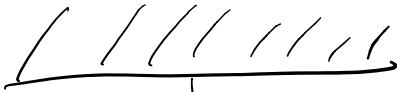
$$= m_1 g \frac{4m_1 + m_2}{4m_1 + m_2}$$

$$T = 3 \frac{m_1 m_2}{4m_1 + m_2} g$$

$$[T] = \frac{kg \cdot kg}{kg} m/s^2 = N \checkmark$$

If  $m_1 = 2m_2$   $a_1 = a_2 = 0 \checkmark$

$$T = ? \quad T = m_1 g$$



$$\begin{aligned} m_1 &= 1 \text{ kg} \\ m_2 &= 2 \text{ kg} \\ m_3 &= 3 \text{ kg} \end{aligned}$$

$$ma = 0$$

$$2T_1 = T_2$$

Find the accelerations of  $m_1$

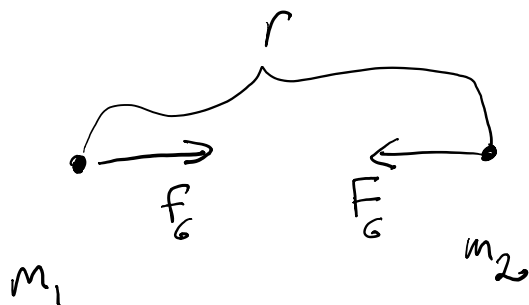
directions  
magnitudes

$$m_1 a_1 = m_1 g - T_1$$

$$m_2 a_2 = m_2 g - T_1$$

$$m_3 a_3 = m_3 g - 2T$$

# Gravitational force

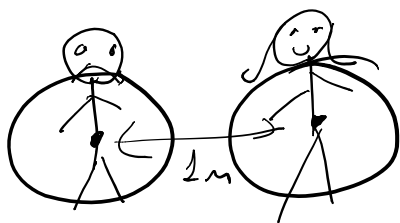


$$F_G = G \frac{m_1 m_2}{r^2}$$

$$F = mg$$

Ex Calculate the force of attraction between Ayşe ( $m = 60 \text{ kg}$ ) and Ali ( $m = 70 \text{ kg}$ )

$$r = 1 \text{ m}$$



$$G = 6.67 \cdot 10^{-11} \frac{\text{Nm}^2}{\text{kg}^2}$$

$$\frac{\text{kg} \cdot \text{kg} \cdot \text{m}^2}{\text{kg}^2}$$

$$\frac{\text{m}^2}{\text{kg}^2}$$

$$F_G = G \frac{m_1 m_2}{r^2} = 6.67 \cdot 10^{-11} \frac{5070}{1^2}$$

$$\approx 33.7 \cdot 10^{-9} \text{ N}$$

$$\approx 3.4 \cdot 10^{-10} \text{ N}$$

$$W_A \approx m_2 g \approx 500 \text{ N}$$

$$F = G \frac{m_1 m_2}{r^2}$$



Motion on earth's surface

$$h \sim 10, 100, 1000 \text{ m}$$

$$F_G = G \frac{m_E m}{(R_E + h)^2}$$

$$R_E \approx 6000 \text{ km}$$

$$\frac{h}{R_E} \sim \frac{1000 \text{ m}}{6000000 \text{ m}} \sim 10^{-4}$$

$$x = 0.01$$

For small values of  $x \ll 1$

$$(1+x)^n = 1 + n \binom{n-1}{1} x + \binom{n}{2} \binom{n-2}{2} x^2 + \dots$$

$\downarrow$  0.01                      0.0001

$$(1+x)^n \approx 1 + nx$$

$$x = 0.0003$$

$$(1+x)^2 = (1.0008)^2 \simeq 1.0006$$

$$F_G = G \frac{m_E m}{(R_E + h)^2} = G \frac{m_E m}{R_E^2 \left(1 + \frac{h}{R_E}\right)^2}$$

$$= G \frac{m_E m}{R_E^2} \left(1 - 2 \frac{h}{R_E}\right)$$

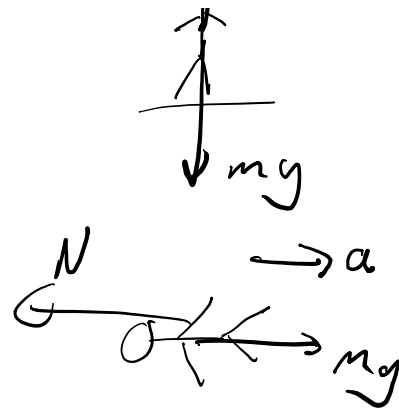
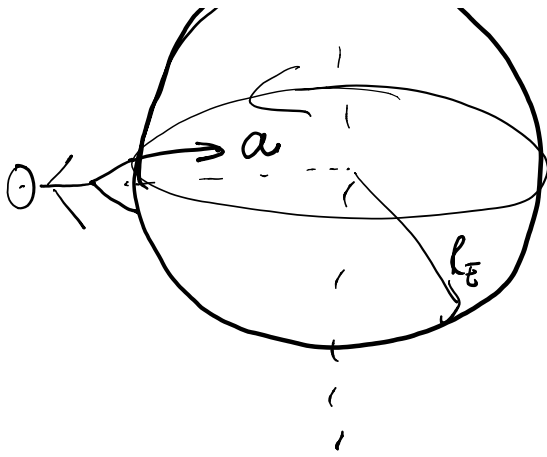
$$F_G = \underbrace{\left(G \frac{m_E}{R_E^2}\right)}_g m - \left(G \frac{m_E}{R_E^2}\right) 2 \frac{h}{R_E} \dots$$

g

$$g = G \frac{m_E}{R_E^2}$$

Ex Find how fast the earth would need to rotate so that the masses on the equator feel weightless!





$$a = \frac{v^3}{R_E} = \omega^2 R_E$$

$$ma = mg$$

$$\omega^2 R_E = g$$

Weightless  $N=0$

$$\omega = \sqrt{\frac{g}{R_E}}$$

$$\omega \approx \sqrt{\frac{10 \text{ m/s}^2}{6 \cdot 10^{25} \text{ m}}} = \sqrt{\frac{1}{6} \cdot 10^{-5}}$$

$$= \sqrt{1.67 \cdot 10^{-6}} = \sqrt{1.67} \cdot 10^{-3}$$

$$= 1.3 \cdot 10^{-3} \frac{1}{\text{s}}$$

Current  $\omega = ?$

$$T = \frac{2\pi}{\omega} = 24 \times 60 \times 60 \text{ sec}$$

$$\omega_{\text{cur}} = 7.27 \cdot 10^{-5}$$

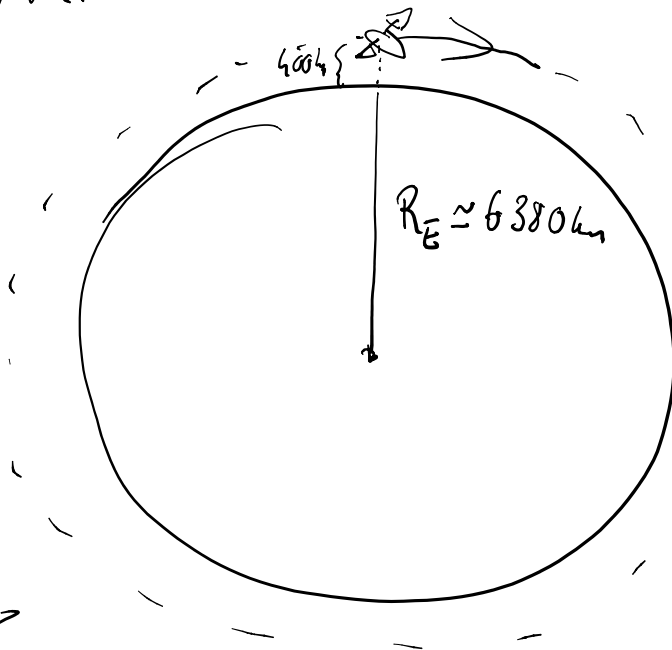




ISS

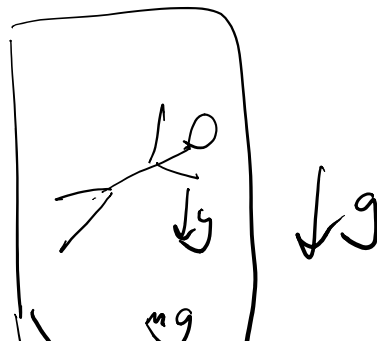
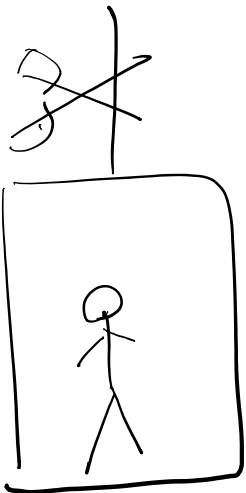
weightless  
astronaut!

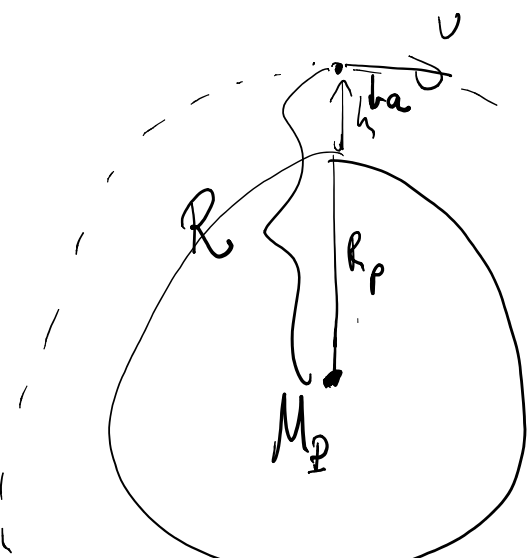
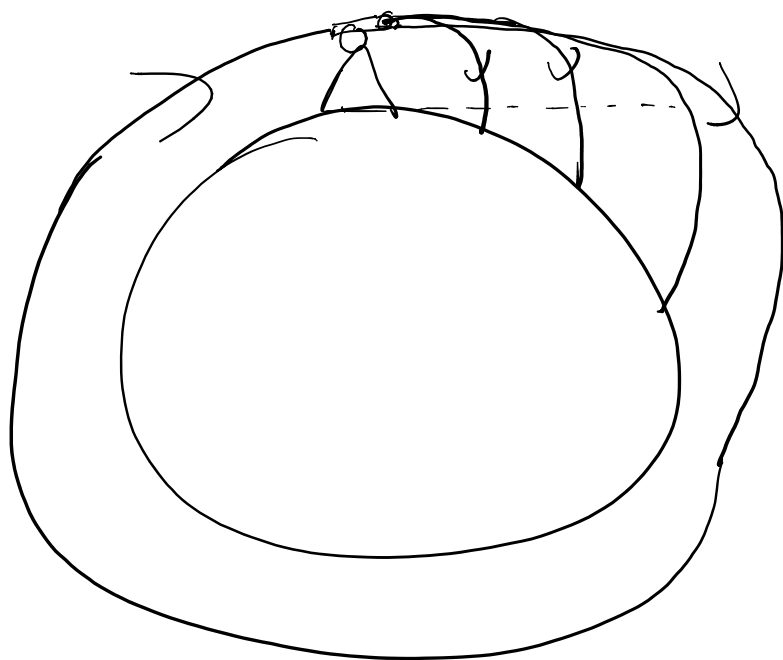
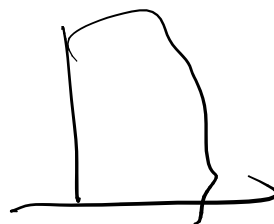
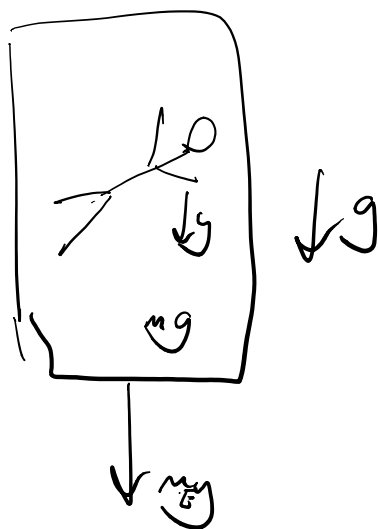
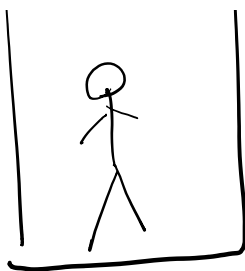
The ISS is about 400 km above  
earth.



$g$  @ ISS

$g \neq 9.82$   
 $\approx 9.8 \text{ m/s}^2$





$$a = \frac{v^2}{R} \downarrow \downarrow F_6$$

$$ma = F_6$$

$$\mu \frac{v^2}{R} = 6 \frac{M_p}{R^2}$$

$$\frac{m}{R} = 6 \frac{m}{R^2}$$

$$V = \sqrt{G \frac{M_p}{R}}$$

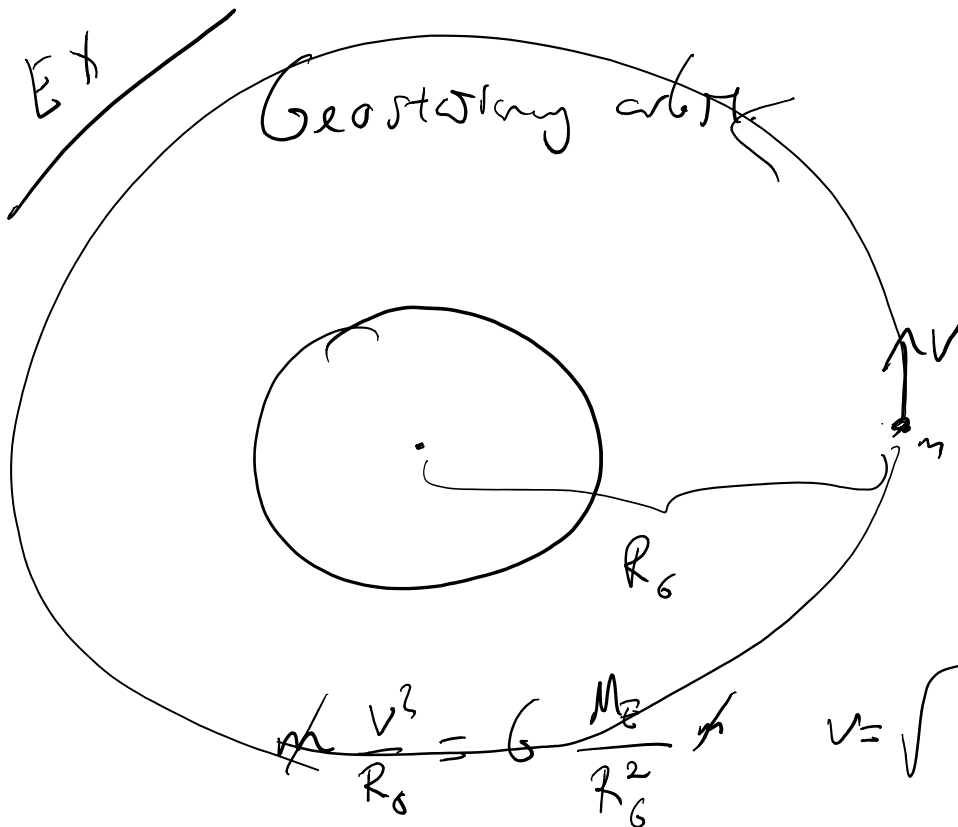
$$G \frac{M_E}{R_E^2} = g$$

$$\approx \sqrt{g R_E}$$

$$V \approx \sqrt{10 \frac{m}{s^2} 6 \cdot 10^6 m}$$

$$\approx \sqrt{10^8 m^2/s^2} = 10^4 m/s$$

$$v_{or} = 10 \text{ km/s}$$



$$T = 24 \text{ h} = \frac{2\pi R_G}{v}$$

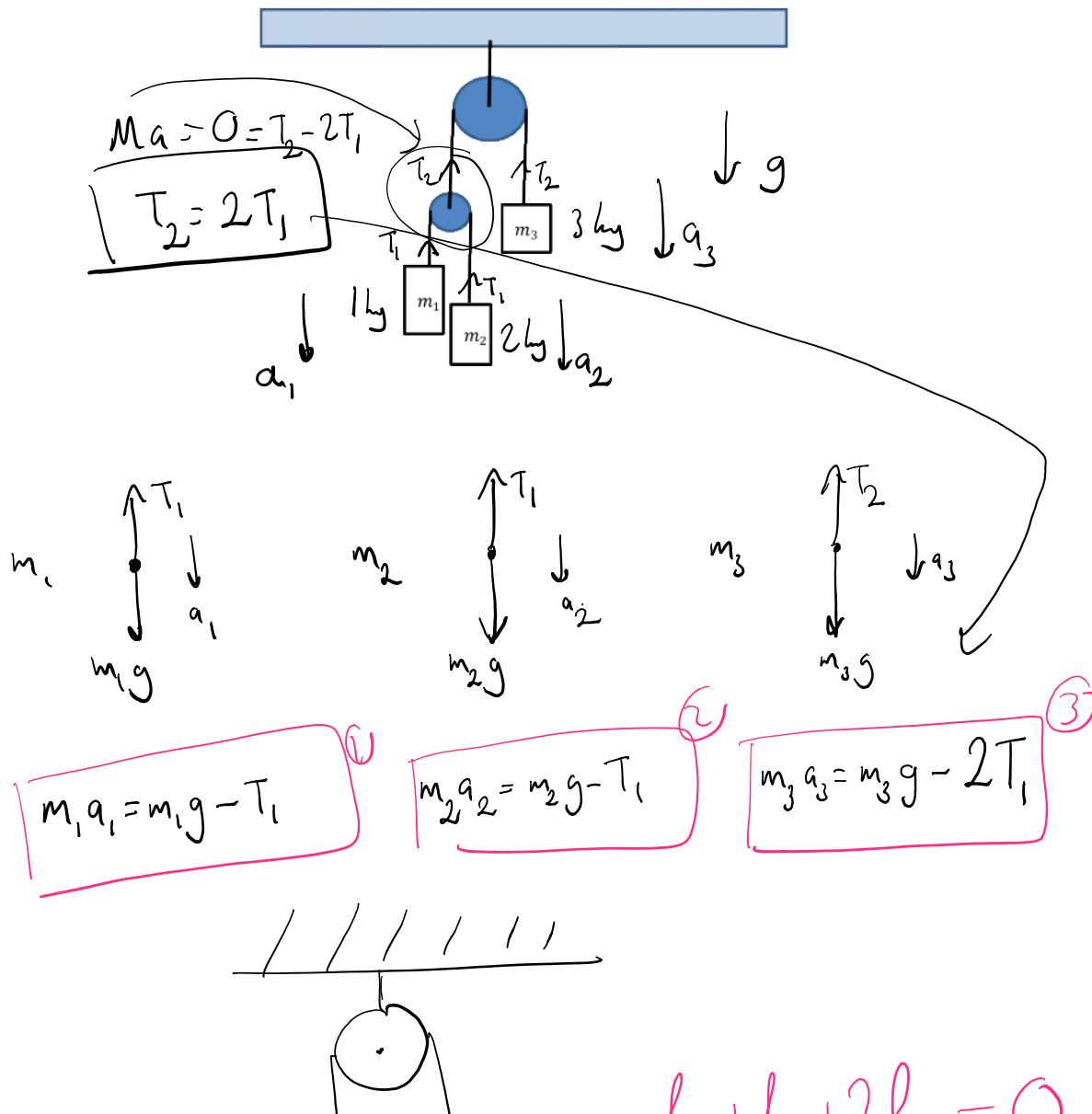
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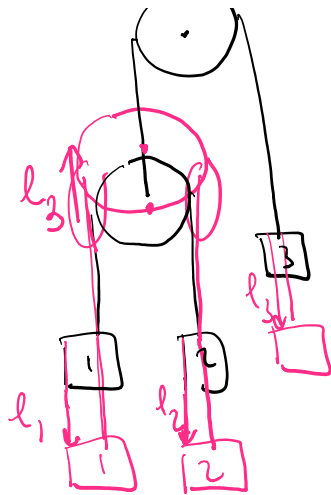
$$24 \times 60 \times 60 = 2\pi \frac{R_6}{\sqrt{G \frac{M_E}{R_6}}} = 2\pi \frac{1}{\sqrt{G M_E}} R_6^{3/2}$$

$$R_6 = \dots$$

The three masses in the figure are  $m_1 = 1.0 \text{ kg}$ ,  $m_2 = 2.0 \text{ kg}$  and  $m_3 = 3.0 \text{ kg}$ . The acceleration due to gravity is given as  $g = 10 \text{ m/s}^2$ . There is no friction, the pulleys and the ropes are massless.

Find the direction and the magnitude of the acceleration of  $m_3$ .





$$l_1 + l_2 + 2l_3 = 0$$

$$\frac{d^2}{dt^2} \left( \begin{matrix} l_1 \\ l_2 \\ 2l_3 \end{matrix} \right) = 0 \quad (4)$$

$$a_1 + a_2 + 2a_3 = 0 \quad (4)$$

$$(a_1 + a_3) = -(a_2 + a_1)$$

$$m_1 a_1 = m_1 g - T \quad (1)$$

$$m_2 a_2 = m_2 g - T \quad (2)$$

$$m_3 a_3 = m_3 g - 2T \quad (3)$$

$$a_1 = g - T$$

$$2a_2 = 2g - T \Rightarrow a_2 = g - \frac{T}{2}$$

$$3a_3 = 3g - 2T \Rightarrow a_3 = g - \frac{2}{3}T \Rightarrow 2a_3 = 2g - \frac{4}{3}T$$

$$a_1 + a_2 + 2a_3 = g - T + g - \frac{T}{2} + 2g - \frac{4}{3}T$$

$$\text{by } (4) \quad 0 = 4g - T \left( 1 + \frac{1}{2} + \frac{4}{3} \right)$$

$$4g = T \left( \frac{\frac{3}{2} + \frac{4}{3}}{\frac{17}{6}} \right) \Rightarrow 4g = T \frac{17}{6}$$

$$T = \frac{24}{17} g$$

$$T = \frac{240}{17} \text{ N}$$

$$m_3 a_3 = m_3 g - 2T$$

$$a_3 = g - \frac{2}{3} T = g - \frac{2}{3} \frac{24}{17} g$$

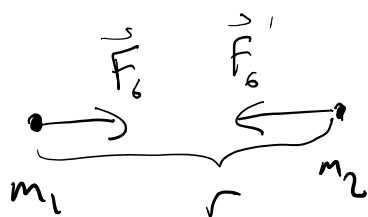
$$10$$

8- 1 3 5 17 ~

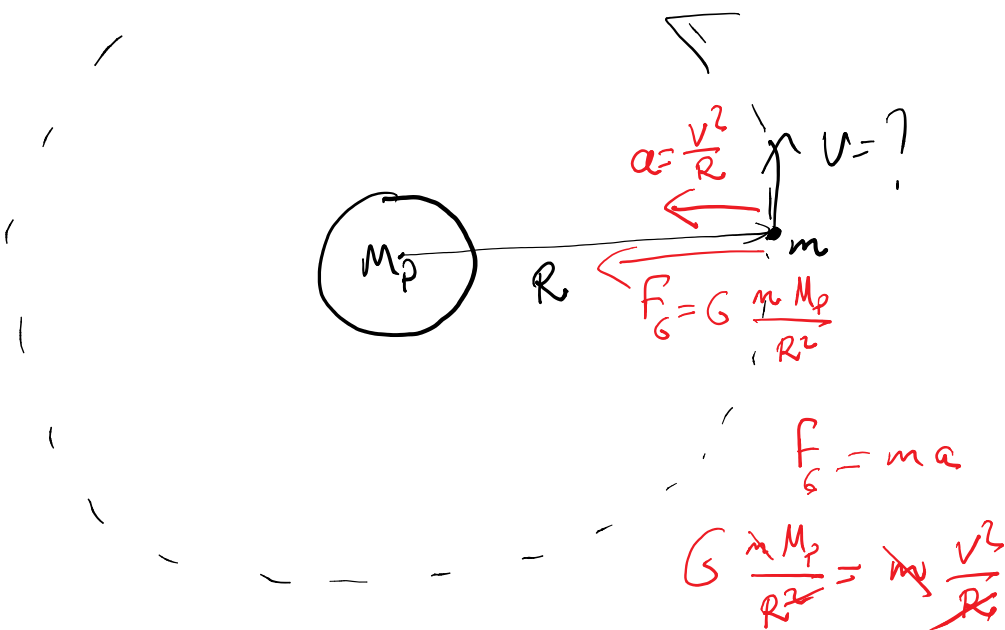
$$a_3 = g \left(1 - \frac{16}{17}\right) = \frac{1}{17} g = \boxed{\frac{10}{17} \text{ m/s}^2}$$

going down!

## Gravitation



$$\boxed{F_g = G \frac{m_1 m_2}{r^2}}$$



$$F_g = m a$$

$$G \frac{M_p}{R^2} = \frac{v^2}{R}$$

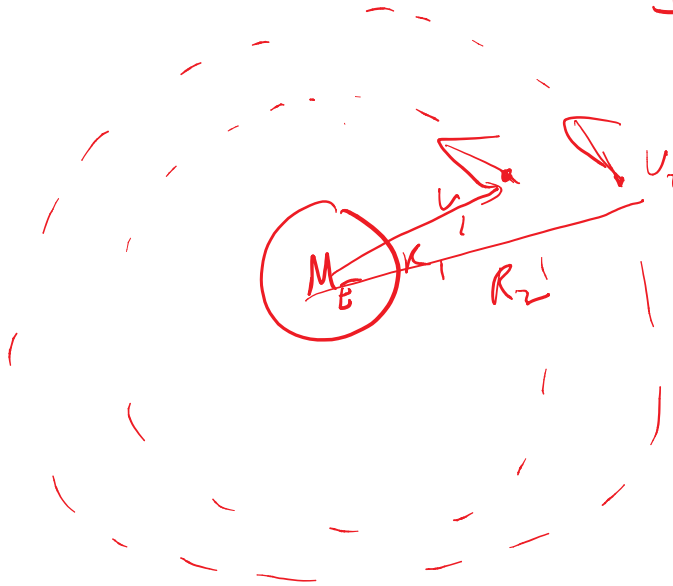
$$v = \sqrt{G \frac{M_p}{R}}$$

What is the period of orbital motion?

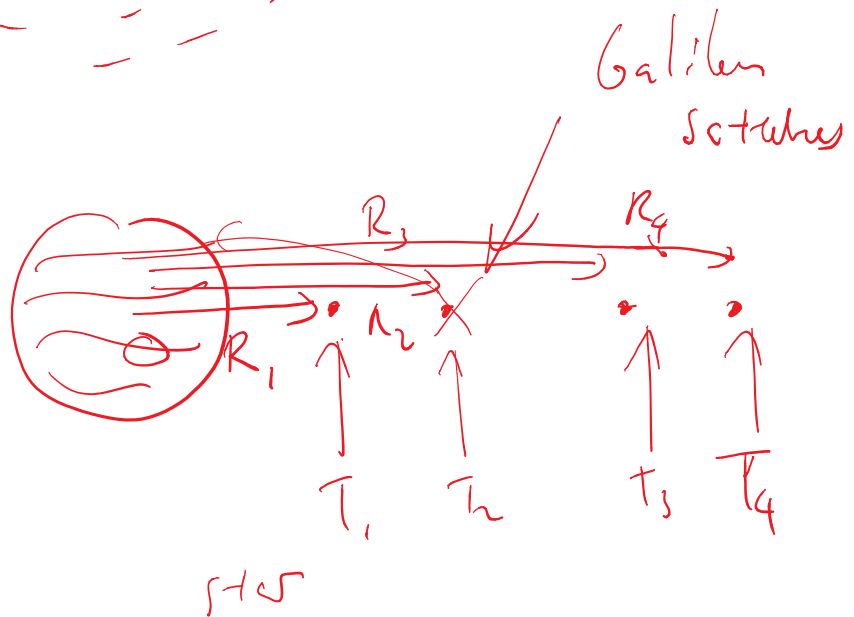
$$T = \frac{2\pi R}{v} = \frac{2\pi R}{\sqrt{G \frac{M_p}{R}}} = 2\pi \frac{R^{3/2}}{\sqrt{G M_p}}$$

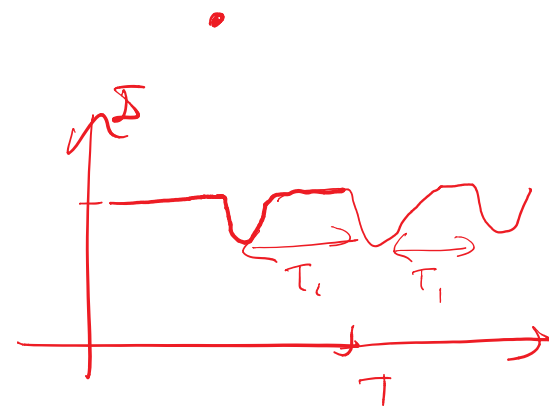
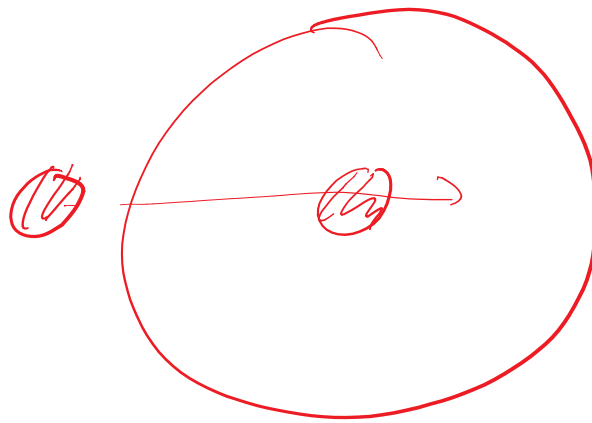
$$v = \frac{2\pi R}{T} = \frac{2\pi R}{\sqrt{\frac{4\pi^2 R^3}{GM_E}}} = \sqrt{\frac{GM_E}{R}}$$

$$T^2 = \frac{4\pi^2 R^3}{GM_E} \Rightarrow \boxed{\frac{R^3}{T^2} = \frac{GM_E}{4\pi^2}}$$



$$\frac{R_1^3}{T_1^2} = \frac{R_2^3}{T_2^2}$$

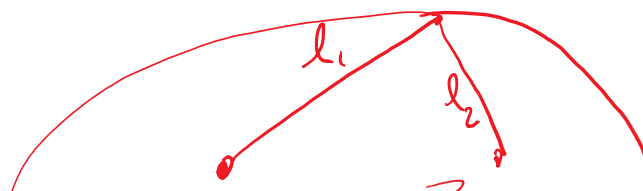
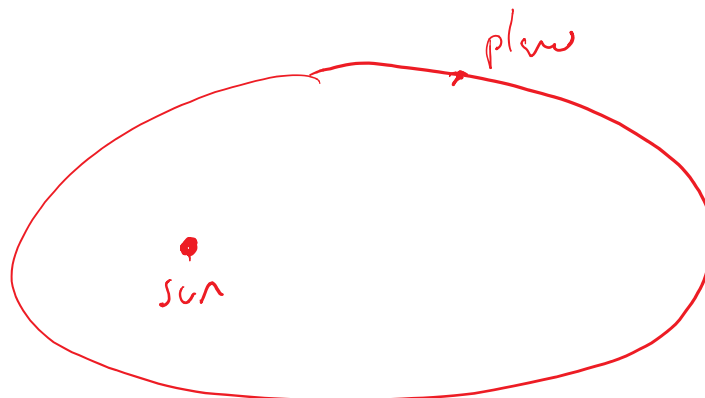




## Kepler's Laws (Tycho Brae)

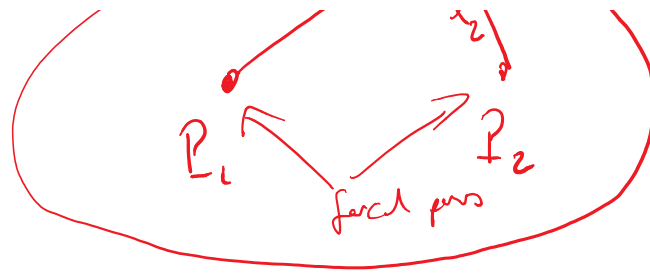
2<sup>nd</sup> Law  $\frac{R^3}{T^2}$  is the same for all planets

1<sup>st</sup> Law Orbit of a satellite is an ellipse with sun at one of the foci.

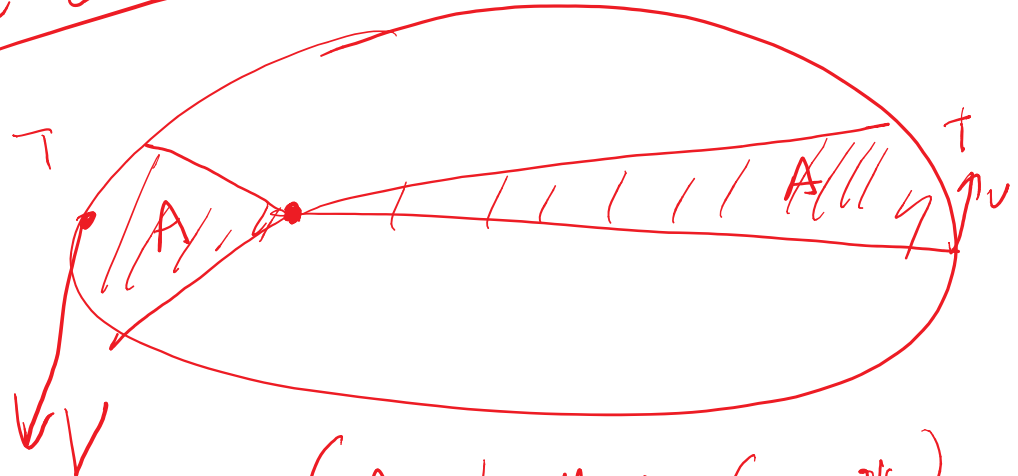


$$l_1 + l_2 = \text{const}$$





2nd Law



(Angular Momentum Conservation)